EIC Accelerator New Technologies and Challenges

BROOKHAVEN +

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> > ENERGY Office of Science

EIC: Electron-Ion Collider

Jefferson Lab

Outline

- The EIC accelerator
 - Requirements and present design
- Accelerator technology challenges
- Some project technology R&D
- Luminosity limiting factors
- Luminosity measurement
- Collider luminosity experience



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Electron-Ion Collider

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"Not Like The Others" (LHC, FCC...)

- The electron-ion collider (EIC) is:
 - a nuclear physics (NP) collider
 - nuclear physics scattering experiments
 - includes "deep inelastic" scattering, or EM-intermediated scattering of electrons and partons
 - collective and single-particle effects in the strong interaction sector
 - NOT a high-energy physics collider
- Addresses three fundamental nuclear physics questions:
 - How does nuclear mass arise?
 - How does nuclear spin arise?
 - What are emergent properties of dense gluon systems?





Ultimately, nuclear tomography 3 Electron-Ion Collider

EIC Requirements

• EIC design goals

- High luminosity: L = (0.1-1)·10³⁴ cm⁻² s⁻¹
 - → 10-100 fb⁻¹
- Collisions of highly polarized (>70%) e and p (and light ion) beams
 - $\,\circ\,$ with flexible bunch by bunch spin patterns
- Large range of CM energies:

○ E_{cm} = 20-140 GeV

- $_{\odot}$ Large range of ion species:
 - Protons Uranium
- Ensure accommodation of a second IR
- Large detector acceptances; good background
 - Hadron particle loss

(IIIII)

IR synchrotron radiation backgrounds

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EIC Requirements

• EIC design goals

- High luminosity: L = (0.1-1)·10³⁴ cm⁻² s⁻¹
 - → 10-100 fb⁻¹ "High"
- Collisions of highly polarized (>70%) e and p (and light ion) beams
 Unique

"Low"

- with flexible bunch by bunch spin patterns
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IR synchrotron radiation backgrounds

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uniquely

challenging



EIC Accelerator Design Overview

Hadron storage ring (HSR): 40-275 GeV (existing)

- up to 1160 bunches, 1A beam current (3x RHIC)
- bright vertical beam emittance (1.5 nm); new vac sleeves
- strong cooling (coherent electron cooling, ERL)

• Electron storage ring (ESR): 2.5–18 GeV (new)

- up to 1160 polarized bunches
 - o high polarization by continual reinjection from RCS
- \circ large beam current (2.5 A) → 9 MW SR power
- o superconducting RF cavities
- Rapid cycling synchrotron (RCS): 0.4-18 GeV (new)
 - 2 bunches at 1 Hz; spin transparent due to high periodicity
- High luminosity interaction region(s) (new)
 - \circ L = 10³⁴ cm⁻²s⁻¹ = 10 kHz-uba, superconducting magnets
 - o 25 mrad crossing angle with crab cavities
 - spin rotators (produce longitudinal spin at IP)

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Luminosity (Lumi) Limits In One Slide™

$L \propto f_{\rm coll} N_1 N_2 / \sigma_x^{\star} \sigma_y^{\star}$

 f_{coll} : collision frequency $N_{1,2}$: particles per bunch $\sigma_{x,y}^{\star}$: (equal) beam sizes at IP

Every parameter optimized separately and collectively in the EIC design

Try multiplying out the given numbers – should be very close to 10³⁴ cm⁻²s⁻¹

- Maximize collision frequency (~90 MHz)
 - Limited by kicker rise times
 - Limited by parasitic collisions (crabbing)
- Maximize particles per bunch (~10¹¹)
 - Limited by sources, space charge
 - Limited by collective effects
 - Interaction of beam with impedances
 - Also total currents: $I_{1,2}=q_{1,2}N_{1,2}f_{coll} \sim 1-3A$
- Minimize beam sizes at IP (~250/25 um)
 - Limited by IR focusing, magnets
 - Limited by chromatic dynamic aperture
 - Limited by emittance growth (IBS)

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Electron-Ion Collider

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Species	proton	electron									
Energy [GeV]	275	18	275	10	100	10	100	5	41	5	6
CM energy [GeV]	140.7		104.9		63.2		44.7		28.6		2
Bunch intensity [10 ¹⁰]	19.1	6.2	6.9	17.2	6.9	17.2	4.8	17.2	2.6	13.3	
No. of bunches	290		1160		1160		1160		1160		_
Beam current [A]	0.69	0.227	1	2.5	1	2.5	0.69	2.5	0.38	1.93	
RMS norm. emit., h/v [μm]	5.2/0.47	845/71	3.3/0.3	391/26	3.2/0.29	391/26	2.7/0.25	196/18	1.9/0.45	196/34	
RMS emittance, h/v [nm]	18/1.6	24/2.0	11.3/1.0	20/1.3	30/2.7	20/1.3	26/2.3	20/1.8	44/10	20/3.5	_
β*, h/v [cm]]	80/7.1	59/5.7	80/7.2	45/5.6	63/5.7	96/12	61/5.5	78/7.1	90/7.1	196/21.0	
IP RMS beam size, h/v [μm]	119	/11	95/8.5		138/12		125/11		198/27		-
K_x	11.1		11.1		11.1		11.1		7.3		
RMS $\Delta \theta$, h/v [µrad]	150/150	202/187	119/119	211/152	220/220	145/105	206/206	160/160	220/380	101/129	_
BB parameter, $h/v [10^{-3}]$	3/3	93/100	12/12	72/100	12/12	72/100	14/14	100/100	15/9	53/42	
RMS long. emittance $[10^{-3}, eV \cdot s]$	36		36		21		21		11		-
RMS bunch length [cm]	6	0.9	6	0.7	7	0.7	7	0.7	7.5	0.7	
RMS $\Delta p / p [10^{-4}]$	6.8	10.9	6.8	5.8	9.7	5.8	9.7	6.8	10.3	6.8	
Max. space charge	0.007	neglig.	0.004	neglig.	0.026	neglig.	0.021	neglig.	0.05	neglig.	
Piwinski angle [rad]	6.3	2.1	7.9	2.4	6.3	1.8	7.0	2.0	4.2	1.1	
Long. IBS time [h]	2.0		2.9		2.5		3.1		3.8		
Transv. IBS time [h]	2.0		2		2.0/4.0		2.0/4.0		3.4/2.1		•
Hourglass factor <i>H</i> 0.91		0.94		0.90		0.88		0.93		. 11: -1	
Luminosity $[10^{33} \text{cm}^{-2} \text{s}^{-1}]$	1.	54	10.00		4.48		3.68		0.44		ollid

Table 3.3: EIC beam parameters for different center-of-mass energies \sqrt{s} , with strong hadron cooling. High divergence configuration.

EIC Accelerator Technology Challenges, R&D



EIC Accelerator Technology Challenges, R&D



HSR Vacuum Cu/aC Coated Beam Screen

- The resistive losses of the 1A proton beam current leads to an unacceptable cryogenic load
 - → need a Cu coated surface
- Build up of electron cloud by ionization of rest gas amplified by secondary emission from stainless steel beam pipe
- Need increase conductivity of RHIC cold SS beampipe, suppress electron cloud with SEY ~1

→ In situ insertion of a Cu and aC (amorphous Carbon) coated screen

• Actively gas-cooled screen is the only feasible solution



EIC Accelerator Technology Challenges, R&D







Focusing (quads) as close to IP as possible (~5m!)

Tensions in magnet requirements:

- high field
- large apertures
- e/p magnet proximity near IP

Chromaticity: focusing dependence on particle energy

EIC R&D: IR Superconducting Magnets



EIC Accelerator Technology Challenges, R&D



EIC R&D: Crab Cavities

- Hadron collider crabbing unprecedented
 - Collaborating with HL-LHC
 - Beam dynamics, RF control stability
 - SPS tests: Phys. Rev. Accel. Beams 24, 062001 (2021)
 - Electron crabbing was performed at KEKB (arXiv:1410.4036)
- Superconducting "RF dipole" cavity
 - ODU design
 - High electric field and overall field quality requirements

• 197 MHz HSR crab cavity being prototyped

- Jefferson Lab/ODU/(BNL) collaboration
- Two possible HOM damping schemes: Waveguide loaded and coaxial couplers
- Stress analysis near completion, with stiffeners



197 MHz Prototype Crab Cavity

Nb

NbTi

Ti

SS

Weld



- Cavity body is comprised of 4 mm Nb with some regions thicker than 4 mm
- Stiffeners are needed to maintain stress at acceptable level



Mechanical Analysis for 197 MHz Crab Cavity

- Cavity body is comprised of 4 mm Nb with some regions thicker than 4 mm
- Stress analysis
 - For VTA test at 22 psi is within allowable stress of 6.3 ksi
- Tuning analysis Tuning in the magnetic field region
 - $\Delta f = \pm 682.3 \text{ kHz}$ (Requirement: $\pm 472 \text{ kHz}$)
 - Tuning sensitivity = 126.4 kHz/mm for a total 5.4 mm displacement
 - 2.7 mm push/pull tuning limit at allowable stress
 - 7400 lb force on each tuner pad (2740 lb/mm)





Prototype Cavity Fabrication

- Majority of the cavity body is deep drawn from 4.17 mm Nb sheets
 - A full Cu cavity will be fabricated to understand and verify the fabrication process
 - Also, to perform room temp measurements
- Flanges:
 - Dogbone flanges Nb with Indium seals
 - All other flanges 316 LN SS CF flanges with Cu gaskets
- All Nb purchased
- JLab actively pressing parts and testing welding approaches

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Electron-Ion Collider

Electrons and Hadrons in Synchrotrons (EIC)

Electrons

Larger charge/mass ratio

- Smaller B to bend/focus (E, crab)
- Normal conducting magnets
- Polarization time dependence

Synchrotron radiation

- Photonic backgrounds
- Damping
- Dynamic aperture: Touschek
- Large RF power needs
- Flat beam aspect ratio
- Harder collimation (multi-stage)

Hadrons

- Smaller charge/mass ratio
 - Larger B to bend/focus (E, crab)
 - Superconducting magnets
 - No depolarization (in principle)

No synchrotron radiation

- Hadronic backgrounds
- Negligible damping (EIC energies)
- Dynamic aperture: "Diffusion"
- Modest RF power needs

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- Round beam aspect ratio
- Easier collimation (single-stage)

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Electron-Ion Collider

Lumi Limits In (more than) One Slide^{NoTM}

$L \propto f_{\rm coll} N_1 N_2 / \sigma_x^{\star} \sigma_y^{\star}$

 f_{coll} : collision frequency $N_{1,2}$: particles per bunch $\sigma_{x,y}^{\star}$: (equal) beam sizes at IP

Challenge: colliding asymmetric beams

electrons: flat hadrons: round SuperKEKB: 10 um x 50 nm, 200:1! EIC collision point: 11:1 transverse aspect ratio U204 JUAS Colliders Session / T. Satogata

Maximize collision frequency (~90 MHz)

- Limited by kicker rise times
- Limited by parasitic collisions (crabbing)

• Maximize particles per bunch (~10¹¹)

- Limited by sources, space charge
- Limited by collective effects
 - Interaction of beam with impedances
 - Also total currents: $I_{1,2}=q_{1,2}N_{1,2}f_{coll} \sim 1-3A$

Electron-Ion Collider

Minimize beam sizes at IP (~250/25 um)

- Limited by IR focusing, magnets
- Limited by chromatic dynamic aperture
- Limited by emittance growth (IBS)

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EIC CDR (CD-1) **Parameters for E_{cm} and Luminosity**

- 18 GeV E _{Cm} = 20 ectrons	41- 275 GeV -140 GeV	10									
E _{Cm} = 20 ectrons	-140 GeV	10									
lectrons	Protons	10	F								
10.0-1/	Protons	-	Ē	Beam-b	eam limi	ted					
IU Gev	275 GeV	S, ₽	E		-	- 19	caling		\mathbf{i}		
E _{Cm} =	105 GeV	cu	-	0	am siz	e lumi s	-				
nb =1160		033	-		pean						
25 mrad										•	
7·10 ¹¹ e	0.7·10 ¹¹ e	osit	E /						10 MV	/ SR limi	ted
2.5 A	1 A	uin 1									
20 nm	9.5 nm	Lur	Spac	e charge	e limited						
1.2 nm	1.5 nm		5	5	10			10	e [GeV]	18	
43 cm	90 cm	01	41	100	100			275	p [GeV]	275	
5 cm	4 cm	2	20	40	60	80	10	00	120	140	1
0.073	0.014		Center of Mass Energy [GeV]								
0.1	0.007										
1.10 ³⁴	cm ⁻² s ⁻¹	$L = 10^{34}$	⁴ cm ⁻² s	⁻¹ = 10	kHz-ub	a					
	E _{cm} − nb = 25 r 7·10 ¹¹ e 2.5 A 20 nm .2 nm 13 cm 5 cm 0.073 0.1 1·10³⁴	E_{cm} - 105 GeV nb =1160 25 mrad $7.10^{11}e$ $0.7.10^{11}e$ $2.5 A$ 1 A $20 nm$ $9.5 nm$.2 nm 1.5 nm $13 cm$ $90 cm$ $5 cm$ 4 cm 0.073 0.014 0.1 0.007 $1.10^{34} cm^{-2} s^{-1}$	E_{cm} - 105 GeV 0 nb =1160 25 mrad $25 mrad$ 1 $7 \cdot 10^{11}e$ $0.7 \cdot 10^{11}e$ $2.5 A$ 1 A $20 nm$ $9.5 nm$ $.2 nm$ $1.5 nm$ $13 cm$ $90 cm$ $5 cm$ 4 cm 0.073 0.014 0.1 0.007 $1 \cdot 10^{34} cm^{-2} s^{-1}$ $L = 10^{34}$	L_{cm} - 105 GeV nb = 1160 25 mrad 7.10 ¹¹ e 2.5 A 1 A 20 nm 9.5 nm .2 nm 1.5 nm 13 cm 90 cm 5 cm 0.073 0.073 0.014 0.1 0.073 0.014 0.1 2024 JUAS Colliders Session / T. Sate	L_{cm} - 105 GeV nb = 1160 25 mrad 7.10 ¹¹ e 2.5 A 1 A 20 nm 9.5 nm .2 nm 1.5 nm 13 cm 90 cm 5 cm 4 cm 0.073 0.014 0.1 0.007 1.10 ³⁴ cm ⁻² s ⁻¹ L = 10 ³⁴ cm ⁻² s ⁻¹ = 10	L_{Cm} - 105 GeV nb = 1160 0.1100 25 mrad 0.7 \cdot 10^{11}e 2.5 A 1 A 20 nm 9.5 nm .2 nm 1.5 nm 43 cm 90 cm 5 cm 4 cm 0.073 0.014 0.1 0.007 1.10 ³⁴ cm ⁻² s ⁻¹ L = 10 ³⁴ cm ⁻² s ⁻¹ = 10 kHz-ub	$\frac{1}{100^{-1}} \frac{100^{-1}}{100^{-1}} 100$	$\frac{E_{cm} - 105 \text{ GeV}}{\text{nb} = 1160}$ $\frac{25 \text{ mrad}}{25 \text{ mrad}}$ $\frac{7 \cdot 10^{11} \text{e}}{2.5 \text{ A}} \qquad 1 \text{ A}$ $\frac{20 \text{ nm}}{9.5 \text{ nm}}$ $\frac{2 \text{ nm}}{1.5 \text{ nm}}$ $\frac{1.5 \text{ nm}}{13 \text{ cm}} \qquad 90 \text{ cm}$ $\frac{5 \text{ cm}}{5 \text{ cm}} \qquad 4 \text{ cm}}{0.073} \qquad 0.014}$ $0.1 \qquad 0.007$ $1 \cdot 10^{34} \text{ cm}^2 \text{s}^{-1}$ $2024 \text{ JUAS Colliders Session / T. Satogata}$ 22	$\frac{1}{100} \frac{1}{25 \text{ mrad}}$ $\frac{7 \cdot 10^{11} \text{e}}{25 \text{ mrad}}$ $\frac{7 \cdot 10^{11} \text{e}}{25 \text{ mrad}}$ $\frac{7 \cdot 10^{11} \text{e}}{25 \text{ A}}$ $\frac{1 \text{ A}}{20 \text{ nm}}$ $\frac{9.5 \text{ nm}}{2.5 \text{ M}}$ $\frac{1 \text{ A}}{20 \text{ nm}}$ $\frac{9.5 \text{ nm}}{1.5 \text{ nm}}$ $\frac{1.5 \text{ nm}}{13 \text{ cm}}$ $\frac{90 \text{ cm}}{5 \text{ cm}}$ $\frac{4 \text{ cm}}{0.073}$ $\frac{0.014}{0.1}$ $\frac{0.1}{0.007}$ $1 \cdot 10^{34} \text{ cm}^2 \text{s}^{-1}$ $2024 \text{ JUAS Colliders Session / T. Satogata}$ 22 Electro	$\frac{E_{Cm} - 103 \text{ GeV}}{\text{nb} = 1160}$ $\frac{25 \text{ mrad}}{25 \text{ mrad}}$ $\frac{7 \cdot 10^{11} \text{e}}{25 \text{ A}} \frac{1 \text{ A}}{1 \text{ A}}$ $\frac{20 \text{ nm}}{9.5 \text{ nm}}$ $\frac{2 \text{ nm}}{1.5 \text{ nm}}$ $\frac{1.5 \text{ nm}}{43 \text{ cm}} \frac{90 \text{ cm}}{5 \text{ cm}} \frac{4 \text{ cm}}{4 \text{ cm}}$ $\frac{0.073}{0.014}$ $\frac{0.073}{0.11} \frac{0.007}{1.10^{34} \text{ cm}^2 \text{s}^{-1}}$ $\frac{2024 \text{ JUAS Colliders Session / T. Satogata}$ $\frac{22}{\text{Electron-lon}}$	$\frac{1}{100^{-1}} \frac{1000}{25 \text{ mrad}}$ $\frac{1}{25 \text{ mrad}}$ $\frac{1}{20 \text{ mr}}$ $\frac{1}$

Lumi Limitations: Space-Charge (low E_{cm})

- Dense charged particle bunches electrostatically repel in rest frame
- Creates **nonlinear** space charge force and equation of motion in lab frame
- Fortunately scales with 1/γ³ so worst at low energies
 - Great example of time dilation

HUM

- Limits high-intensity injector emittances
- Force applies continuously within beam
- Tolerable linear "space charge tune spread" of 0.05 limits total current of 41 GeV proton beam to ~0.4A.

(IBS: intra-beam hard scattering also contributes)



Lumi Limitations: Beam-Beam (mid E_{cm})

- Colliding beams see each other's collective charge distributions
- Creates nonlinear beam-beam force and equation of motion similar to space charge
 - **BUT** now the fields and force are in the lab frame already
 - NO benefit of relativistic scaling
 - Force applies only once per turn
- Tolerable "beam-beam tune spead" of 0.015 limits highest EIC luminosity

$$F(r) = \frac{Nq^2}{2\pi\epsilon_0 l} \frac{1+\beta^2}{r} \left[1 - \exp\left(-\frac{r^2}{2\sigma^2}\right)\right]$$



Lumi Limitations: Electron SR Power (high E_{CM})

- Accelerated charged particles emit photons
 - Electrons in synchrotron: radially accelerated
 - Synchrotron radiation emitted in forward cone
 - Cone opening angle $\propto 1/\gamma$
 - Radiated power $P_{\gamma} = \frac{2}{3} \frac{e^2 c}{4\pi\epsilon_0} \frac{(\gamma\beta)^4}{\rho^2}$
 - γ scaling **much** worse for electrons
 - 18 GeV e: γ =3.5x10⁴ vs 255 GeV p: γ =3x10²
- Design: 9 MW @ 18 GeV (facility limit 10 MW)
- Expensive: Power must be provided by SRF
- Raise electron energy last (e current limit)



Collider Luminosity Ramp-Up

- Luminosity ramp-up to design takes years
 - Useful paper: arxiv 1202.3950 (V. Shiltsev)
 - Contextualizes Tevatron Run-II and early LHC
 - Luminosity ramp-up parameter C: complexity
 - C: time (years) to increase lumi by e
 - C=2: factor of e luminosity increase in 2 years
 - Early commissioning can make quick strides
 - C<1 (or <<1) but do not get too exuberant
 - Long-term commissioning usually C~2-3
- EIC will very likely take years to reach design luminosities
 - But we will get there!





Tevatron Run-II: design 275x10³⁰

K. Piotrzkowski, <u>https://arxiv.org/pdf/2106.08993.pdf</u> 2021 *JINST* **16** P09023

EIC Luminosity Measurement

- Bethe-Heitler bremsstrahlung
 - Induced e- radiation
 - Proportional to luminosity
 - With correction terms
 - Very "e-forward" electrons
 - Similar to synchrotron radiation

Challenges

- Bremsstrahlung rate suppression due to the so-called beam size effect (observed at HERA)
- Huge synchrotron radiation fluxes should be mitigated (split dipole)
- Enormous bremsstrahlung event rates, up to 10 GHz



Table 1. Bethe-Heitler *ep* bremsstrahlung cross sections in mb (and the corresponding event rates in GHz, for the nominal EIC luminosities), for various beam energies in GeV and three selection criteria.

		-			
E_e	E_p	$E_{\gamma}/E_e > 0.01$	$1 > E_{\gamma}/E_e > 0.7$	$0.4 > E_{\gamma}/E_e > 0.1$	
18	275	237 (0.36)	11.6 (0.018)	65.2 (0.10)	
10	275	230 (2.3)	11.1 (0.11)	63.2 (0.63)	
5	100	209 (0.77)	9.81 (0.036)	57.1 (0.21)	
		,			
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EIC Primary Interaction Region: Lumi Monitor



Summary

- EIC design meets all design requirements
- EIC luminosity is highly optimized
 - Balances several individual parameter optimizations
- EIC R&D progressing
 - Focus: challenging technical components
- e/p beam differences drive EIC choices
 - e.g. luminosity vs E_{cm} , IP aspect ratio
- Luminosity ramp-up will take time
 - will last substantially beyond project end
 - project provides excellent basis to get there!





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